

One-point structure tensors in Rayleigh-Taylor turbulence

Benoît-Joseph Gréa, Jérome Griffond and Olivier Soulard

CEA/DAM/DIF F-91297 Arpajon, FRANCE

IWPCTM conference - 10-17 July 2010

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Motivatio

DNS of RT during linear phase

Structure tensors

Outline

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Linear phases (RDT) in RT with complex acceleration history



Acceleration modification during RT instability changes automatically the Froude number defined as:

$$\mathcal{F}_r = \frac{\varepsilon}{kN}$$
 with $N = \left| \frac{g}{\overline{\rho}} \frac{\partial \overline{\rho}}{\partial z} \right|^{1/2}$. (1)

When $\mathcal{F}_r \ll 1$ we have a "RDT" phase (Rapid Distorsion theory) *i.e.* non-linear terms are negligeable compared to the buoyancy terms in momentum/concentration equations for turbulent fluctuations.

The classical linear RDT equations for unstable, homogeneous, low Atwood, incompressible flow write in Fourier space:

$$(\partial_t - \nu \kappa^2) \hat{u}_i(\vec{\kappa}, t) = -\left(\delta_{i3} - \frac{\kappa_i \kappa_3}{\kappa^2}\right) \hat{\rho}(\vec{\kappa}, t),$$
(2)

$$(\partial_t - \lambda \kappa^2) \hat{\rho}(\vec{\kappa}, t) = -N^2 \hat{u}_3(\vec{\kappa}, t).$$
(3)

cf. A.Townsend (1976), H.Hanazaki et al. (1995)

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Difficulty of modeling linear phases for single point turbulent models



One-point turbulent models (based on R_{ij} for instance) are deficient when capturing RDT phases. Indeed,

- rapid pressure correlation terms are strongly dependent on spectral wave number directionality;
- the characteristics of turbulence usually change a lot during RDT phases;
- classical one-point turbulent models lack informations. For instance, *R_{ij}* gives "componentality" but says nothing about "dimensionality" of turbulence.

cf. Kassinos et al., JFM 428 (2001)



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Example of linear phase in homogeneous unstable stratification

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Analytic solutions for RDT equations exist and can be compared to classical models.

For example, assuming that Fourier expressions \hat{R}_{ij} , $\hat{\rho}\hat{u}_3$, $\hat{\rho}\hat{\rho}$ depends on κ only (and not $\vec{\kappa}$) one obtains a simple closure:

$$\frac{d < u_3 u_3 >}{dt} = -\frac{4}{3} < \rho u_3 >, \tag{4}$$

$$\frac{d < \rho u_3 >}{dt} = -N^2 < u_3 u_3 > -\frac{2}{3} < \rho \rho >, \tag{5}$$

$$\frac{d < \rho \rho >}{dt} = -2N^2 < \rho u_3 > . \tag{6}$$

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Example of linear phase in homogeneous unstable stratification



As expected, very bad agreement between model and theory.

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Objective



The objective is

- 1. Study RT during RDT phases with DNS
- 2. Define new tensors to take into account dimensionality of $<\rho\rho>$
- 3. Propose a closure for linear phase using the new tensors



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TRICLADE code (cf M.Boulet and J.Griffond 10th IWPCTM)



Solves 3D compressible Navier-Stokes+concentration equations for mixtures of perfect gases.

Presently used scheme (hyperbolic part):

 high (5th) order one-step scheme with uniform time and space accuracy;

cf. V. Daru and C. Tenaud, JCP 193 (2004)

- directional splitting;
- direct Euler solver (not Lagrange+projection);
- wave propagation method;

cf. R.J LeVêque

different Riemann solvers.

cf. E.F. Toro

Presently used scheme (elliptic part+sources): operator splitting, 2nd order treatment for viscous-diffusive terms and for sources.

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Simulation characteristics



- mesh 512 \times 256 \times 256, H = 1m;
- Atwood number At = 0.1, viscosity $\nu = 10^{-5}m^2s^{-1}$, Schmidt number Sc = 1 and Prandtl number $Pr \approx 3/4$;
- initial stratification adiabatic with $\gamma = 5/3$, pressure $p_0 \approx 4m^2s^{-2}$ (incompressible limit);
- Initial perturbations for RT instability " M256 type" ;

cf. G.Dimonte et al., Phys. Fluids 16 (2004)

- Gravity: $G = 0.2ms^{-2}$ for $T \le 14s$ before RDT phase, then $G = 20ms^{-2}$ for T > 14s;
- Rescaling of mean pressure at the beginning of RDT phase to avoid shock creation.

cf. A.Llor and D. Youngs IWPCTM (2002)

Global results



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Determination of linear phase



Linear phase (observed), between T = 14s and T = 14.15s:

- mixing zone length is still, $N = 2.7s^{-1}$;
- turbulent kinetic energy explodes.

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Validity of RDT equations



Validity of RDT equations



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Validity of RDT equations



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Validity of RDT equations



Good agreement at large scale where energy is contained

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Spectra



Structure tensors

Spectra



Small modification of concentration field.

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Spectra





One-point structure tensors in Rayleigh-Taylor during linear RDT phases

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Hydrodynamic tensors from Kassinos and Reynolds

One-point hydrodynamic tensors are defined by:

Componentality
$$R_{ij} = \epsilon_{ipq}\epsilon_{jts} < \psi_{q,p}\psi_{s,t} >_{xy},$$
 (7)

Dimensionality
$$D_{ij} = \langle \psi_{n,i} \psi_{n,j} \rangle_{xy},$$
 (8)

Circulicity
$$F_{ij} = \langle \psi_{i,n} \psi_{j,n} \rangle_{xy},$$
 (9)

"Inhomogeneity"
$$C_{ij} = \langle \psi_{i,n} \psi_{n,j} \rangle_{xy}$$
 (10)

where the stream vector $\vec{\psi}$ is solution of

$$\psi_{i,nn} = -\omega_i, \quad \psi_{i,i} = \mathbf{0}, \quad \mathbf{U}_i = \epsilon_{its} \psi_{s,t}.$$
 (11)

cf. Kassinos et al., JFM (2001) also C.Cambon et al. Phys. Fluids 4 (1992).

They are not independent as,

$$R_{ij} + D_{ij} + F_{ij} - (C_{ij} + C_{ji}) = 2k^2 \delta_{ij}.$$
 (12)

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Evolution of purely kinematic tensors



Evolution of purely kinematic tensors

"Inhomogeneity" tensor is equal to 0:

Poloidal velocity correlations negligeable.

$$\vec{u} = \vec{\nabla} \times \underbrace{\mathbf{s}_{\text{pol}} \mathbf{x}_3}_{(0, 0, \psi_3)} + \vec{\nabla} \times \underbrace{\vec{\nabla} \times \mathbf{s}_{\text{tor}} \mathbf{x}_3}_{(\psi_1, \psi_2, 0)}$$
(13)



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Representation of purely kinematic tensors in Lumley triangle



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Definition of new velocity/tensor

We define an enhanced velocity by adding a solenoidal part dependent on density ρ :

$$\hat{v}_{i}(\vec{\kappa},t) = \hat{u}_{i}(\vec{\kappa},t) + \hat{u}_{i}^{(\rho)}(\vec{\kappa},t), \text{ with } \hat{u}_{i}^{(\rho)}(\vec{\kappa},t) = \frac{\kappa_{i}}{\kappa N} \hat{\rho}(\vec{\kappa},t).$$
(15)

cf. J.J. Riley and al (1981), F.S. Godeferd and C. Cambon Phys Fluid 6 (1994) The time evolution in RDT is determined by the linear equation:

$$\frac{d}{dt}\hat{v}_i = L_{ij}\hat{v}_j,\tag{16}$$

with

$$L_{ij} = -N\left(\mathcal{P}_{i3}\frac{\kappa_j}{\kappa} + \mathcal{P}_{j3}\frac{\kappa_i}{\kappa}\right), \text{ with } \mathcal{P}_{ij} = \delta_{ij} - \frac{\kappa_i \kappa_j}{\kappa^2}.$$
(17)

The second order spectral tensor can be formed from enhanced velocity:

$$\hat{V}_{ij}(\vec{\kappa}) = \langle \hat{v}_i(\vec{\lambda})\hat{v}_j^*(\vec{\kappa}-\vec{\lambda}) \rangle$$
(18)

$$\frac{d}{dt}\hat{V}_{ij} = L_{in}\hat{V}_{nj} + L_{jm}\hat{V}^*_{mi}$$
(19)

Dimensionality for ρ defined as:

$$\hat{D}_{ij}^{(\rho)}(\vec{\kappa}) = \langle u_i^{(\rho)}(\vec{\lambda}) u_j^{(\rho)*}(\vec{\kappa} - \vec{\lambda}) \rangle$$
(20)

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evolutions of new tensor in physical space





Representation of mixed tensors in Lumley triangle



cf. different from dimensionality tensor X.Albets-Chico and al. IUTAM (2010).

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Anisotropy decomposition in linear RT

Representation of the enhanced velocity tensor in a Craya-Herring frame



The Craya-Herring frame is defined by, $e^1 = \vec{\kappa} \times n/|\vec{\kappa} \times n|, e^2 = \vec{\kappa} \times e^1/|\vec{\kappa} \times e^1|, e^3 = \vec{\kappa}/|\vec{\kappa}|.$



We have for axisymetric turbulence,

$$\frac{1}{2} \mathbf{e}^{\mathbf{i}}_{n} \hat{V}_{nm} \mathbf{e}^{\mathbf{j}}_{m} = \begin{pmatrix} \Phi_{1} & 0 & 0 \\ 0 & \Phi_{2} & \Psi^{*} \\ 0 & \Psi & \Phi_{3} \end{pmatrix}$$

where

- Φ₁ represents turbulent kinetic energy of poloidal velocity;
- Φ₂ is the turbulent kinetic energy of toroidal velocity;
- Φ_3 define a potential energy from the density spectrum;
- Ψ generates the vertical buoyancy flux.

$$\frac{1}{2} < u_3 \rho >= N \int \Psi_R \sin(\theta) d^3 \vec{\kappa}$$
(22)

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Anisotropy decomposition

Expansion for the tensors

The evolution equations write:

 $\frac{\partial \Phi_1}{\partial t} = 0, \quad \frac{\partial \Psi_I}{\partial t} = 0, \tag{23}$

$$\frac{\partial \Phi_2}{\partial t} - Nsin(\theta)\Psi_R = 0, \qquad (24)$$

$$\frac{\partial \Phi_3}{\partial t} - N \sin(\theta) \Psi_R = 0, \tag{25}$$

$$\frac{\partial \Psi_R}{\partial t} - 2Nsin(\theta)(\Phi_2 + \Phi_3) = 0.$$
(26)

This suggests to use the decomposition:

$$\frac{\Phi_2(\vec{\kappa},t) + \Phi_3(\vec{\kappa},t)}{2} = H_0^{(dir)}(\kappa,t) + H_1^{(dir)}(\kappa,t)\sin^2\theta + \dots + H_n^{(dir)}(\kappa,t)\sin^{2n}\theta + \dots$$
(27)

$$\Psi_{R}(\vec{\kappa},t) = \sin\theta \left(H_{0}^{(\text{pola})}(\kappa,t) + H_{1}^{(\text{pola})}(\kappa,t)\sin^{2}\theta + \dots + H_{n}^{(\text{pola})}(\kappa,t)\sin^{2}\theta + \dots \right)$$
(28)

Compatibility with constraint of polar isotropy, at $\kappa_1 = \kappa_2 = 0$, $\Psi = 0$ and $\Phi_1 = \Phi_2$.

$$\frac{\partial H_n^{pola}}{\partial t} = 2NH_n^{dir}, \ \frac{\partial H_{n+1}^{dir}}{\partial t} = NH_n^{pola}.$$
 (29)

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Expansion for the tensors

Assuming to simplify that poloidal velocity equals 0 ($\Phi_1 = 0$) we have:

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$$\frac{1}{2}R(t) = \underbrace{\int \frac{\Phi_2 - \Phi_3}{2}(\vec{\kappa})M^{(u)}(\vec{\kappa})d^3\vec{\kappa}}_{=} + \underbrace{\int \frac{\Phi_2 + \Phi_3}{2}(\vec{\kappa}, t)M^{(u)}(\vec{\kappa})d^3\vec{\kappa}}_{=} \qquad (30)$$
$$= R^{\text{polarized}} + \sum_n R^{(n)} \int H_n^{(\text{dir})} d\kappa(t) \qquad (31)$$

in the same way,

$$\frac{1}{2}D_{(\rho)}(t) = D_{(\rho)}^{\text{polarized}} + \sum_{n} D^{(n)} \int H_{n}^{(\text{dir})} d\kappa(t)$$
(32)

with

$$R^{(n)} = \int \kappa_H^n M^{(u)} d^3 \vec{\kappa}, \ D^{(n)} = \int \kappa_H^n M^{(\rho)} d^3 \vec{\kappa}.$$
(33)

$$M^{(\mu)}(\vec{\kappa}) = \begin{pmatrix} \frac{\kappa_1^2 \kappa_3^2}{\kappa_H^2 \kappa^2} & \frac{\kappa_1 \kappa_2 \kappa_3^2}{\kappa_H^2 \kappa^2} & -\frac{\kappa_1 \kappa_3}{\kappa^2} \\ \frac{\kappa_1 \kappa_2 \kappa_3^2}{\kappa_H^2 \kappa^2} & \frac{\kappa_2^2 \kappa_3^2}{\kappa_H^2 \kappa^2} & -\frac{\kappa_2 \kappa_3}{\kappa^2} \\ -\frac{\kappa_1 \kappa_3}{\kappa_H^2 \kappa^2} & \frac{\kappa_2^2 \kappa_3}{\kappa_H^2 \kappa^2} & \frac{\kappa_H^2}{\kappa^2} \end{pmatrix}, M^{(\rho)}(\vec{\kappa}) = \begin{pmatrix} \frac{\kappa_1^2}{\kappa^2} & \frac{\kappa_1 \kappa_2}{\kappa^2} & \frac{\kappa_1 \kappa_3}{\kappa^2} \\ \frac{\kappa_1 \kappa_2}{\kappa^2} & \frac{\kappa_2^2 \kappa_3}{\kappa^2} & \frac{\kappa_2 \kappa_3}{\kappa^2} \\ -\frac{\kappa_1 \kappa_3}{\kappa^2} & -\frac{\kappa_2 \kappa_3}{\kappa^2} & \frac{\kappa_H^2}{\kappa^2} \end{pmatrix}, M^{(\rho)}(\vec{\kappa}) = \begin{pmatrix} \frac{\kappa_1^2}{\kappa^2} & \frac{\kappa_1 \kappa_2}{\kappa^2} & \frac{\kappa_1 \kappa_3}{\kappa^2} \\ \frac{\kappa_1 \kappa_3}{\kappa^2} & \frac{\kappa_2 \kappa_3}{\kappa^2} & \frac{\kappa_3}{\kappa^2} \end{pmatrix} d^3 \vec{\kappa},$$
(34)

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Model for RDT using new tensor

Back to the first example!



Improvement due to better modeling of anisotropy evolution.

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Structure tensors

Conclusion

DNS of linear RDT phase for Rayleigh-Taylor instability	
have been performed.	

- An analysis by one-point structure tensors has been presented.
- An original decomposition of anisotropy has been realized based on Craya-Herring representation of enhanced velocity.
- This allows a new one-point closure for linear RT phase based on dimensionality of $< \rho \rho >$.

Thank you for your attention

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